#### Computer Physics Communications 219 (2017) 223-235

Contents lists available at ScienceDirect

## **Computer Physics Communications**

journal homepage: www.elsevier.com/locate/cpc



# A robust and efficient numerical method to compute the dynamics of the rotating two-component dipolar Bose–Einstein condensates



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#### ARTICLE INFO

Article history: Received 17 June 2016 Received in revised form 12 May 2017 Accepted 23 May 2017 Available online 1 June 2017

Keywords: Two-component dipolar BEC Dynamics Gaussian-sum method Rotating Lagrangian coordinates Time splitting Fourier spectral method Collapse dynamics

#### ABSTRACT

We present a robust and efficient numerical method to compute the dynamics of the rotating twocomponent dipolar Bose–Einstein condensates (BEC). Using the rotating Lagrangian coordinates transform (Bao et al., 2013), we reformulate the original coupled Gross–Pitaevskii equations (CGPE) into new equations where the rotating term vanishes and the potential becomes time-dependent. A time-splitting Fourier pseudospectral method is proposed to numerically solve the new equations where the nonlocal Dipole–Dipole Interactions (DDI) are computed by a newly-developed Gaussian-sum (GauSum) solver (Exl et al., 2016) which helps achieve spectral accuracy in space within  $O(N \log N)$  operations (N is the total number of grid points). The new method is spectrally accurate in space and second order accurate in time — these accuracies are confirmed numerically. Dynamical properties of some physical quantities, including the total mass, energy, center of mass and angular momentum expectation, are presented and confirmed numerically. Interesting dynamical phenomena that are peculiar to the rotating twocomponent dipolar BECs, such as dynamics of center of mass, quantized vortex lattices dynamics and the collapse dynamics in 3D, are presented.

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## 1. Introduction

The Bose-Einstein condensation (BEC) yields most interesting state-of-the-art experiments for relatively large quantum systems and has been extensively studied since its first experimental realization in 1995 [1-3]. A subsequent achievement of quantum vortices in rotating BECs [4-6] broadens the attention to explore vortex states and their dynamics associated with superfluidity. Initially, the experiments/simulations were limited to the case of short-range interatomic interactions [7]. Recently, considerable attention has been drawn to systems with long-range dipoledipole interactions (DDIs) in ultracold physics [8]. For heteronuclear molecules, DDIs come from their electric dipole moments [9]. In a state with a well-defined angular momentum, molecules do not have a dipole moment. Nevertheless, dipolar moments can be induced when molecules are polarized via an external electric field. For atoms, dipolar interactions arise from their magnetic moments and become significant for large electronic spin. Recent experiments on dipolar BECs of Cr atoms and others [10-12] demonstrated very well for such interactions and has spurred new impetus in the study of dipolar BECs.

Due to the presence of anisotropic DDI, vortices in rotating dipolar BECs exhibit novel properties and richer phenomena [13–16]. On the other hand, thanks to the development of trapping techniques, binary condensates are also realized [17–19] and provide an ideal system for studying phase transitions and coexistence of different phases [20–22]. Far from being a trivial extension of the single-component BEC, the physics of a dipolar mixture may feature different and remarkable properties, such as the domain walls, vortons and square vortex lattices [18,19,23].

Very recently, the vortices of rotating two-component dipolar BECs under different trapping potentials have been investigated in several studies by physicists [22–26]. At temperatures T much smaller than the critical temperature  $T_c$ , the properties of rotating two-component dipolar BECs are well described by the macroscopic complex-valued wave function  $\Psi = (\psi_1(\mathbf{x}, t), \psi_2(\mathbf{x}, t))^T$  calculated from the three-dimensional (3D) Coupled Gross-Pitaevskii Equations (CGPE) with DDI term. Moreover, the 3D CGPE can be reduced to an effective two-dimensional (2D) equation if the external potential is highly anisotropic, i.e. much stronger in *z*-direction [27,28]. In a unified way, the *d*-dimensional (*d* = 2 or 3) dimensionless CGPE with DDI term

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reads as [23,25,26,29,30]:

$$i\partial_t \psi_j(\mathbf{x}, t) = \left[ -\frac{1}{2} \nabla^2 + V_j(\mathbf{x}) - \Omega L_z + \sum_{k=1}^2 \left( \beta_{jk} |\psi_k|^2 + \lambda_{jk} \, \Phi_k(\mathbf{x}, t) \right) \right] \psi_j(\mathbf{x}, t), \qquad (1.1)$$

$$\begin{aligned}
\varphi_{j}(\mathbf{x},t) &= U_{\rm dip} * |\psi_{j}|^{2}, \quad \psi_{j}(\mathbf{x},t=0) = \psi_{j}^{\circ}(\mathbf{x}), \\
j &= 1, 2, \quad \mathbf{x} \in \mathbb{R}^{d}, \quad t \ge 0.
\end{aligned} \tag{1.2}$$

Here,  $\Phi_j$  is defined as a convolution of the kernel  $U_{dip}$  with the density  $|\psi_j|^2$  where \* denotes the convolution operator, t denotes time,  $\mathbf{x} = (x, y, z)^T \in \mathbb{R}^3$  and/or  $\mathbf{x} = (x, y)^T \in \mathbb{R}^2$  is the Cartesian coordinate vector. The constant  $\beta_{jk}$  describes the strength of the short-range interactions in a condensate (positive/negative for repulsive/attractive interaction),  $L_z = -i(x\partial_y - y\partial_x) = -i\partial_\theta$  is the z-component of the angular momentum and  $\Omega$  represents the rotating frequency.  $V_j(\mathbf{x})$  (j = 1, 2) is a given real-valued external trapping potential determined by the type of system under investigation. In most BEC experiments, a harmonic potential is chosen to trap the condensates, i.e. for j = 1, 2

$$V_{j}(\mathbf{x}) = \frac{1}{2} \begin{cases} \gamma_{x,j}^{2} x^{2} + \gamma_{y,j}^{2} y^{2}, & d = 2, \\ \gamma_{x,j}^{2} x^{2} + \gamma_{y,j}^{2} y^{2} + \gamma_{z,j}^{2} z^{2}, & d = 3, \end{cases}$$
(1.3)

where  $\gamma_{v,j}$  (v = x, y, z) are dimensionless constants representing the trapping frequencies in *v*-direction. Moreover,  $\lambda_{ij}$  (i, j = 1, 2) is a constant characterizing the strength of DDI and  $U_{dip}(\mathbf{x})$  is the long-range DDI potential. In 3D,  $U_{dip}(\mathbf{x})$  reads as

$$U_{\rm dip}(\mathbf{x}) = \frac{3}{4\pi |\mathbf{x}|^3} \left[ 1 - \frac{3(\mathbf{x} \cdot \mathbf{n})^2}{|\mathbf{x}|^2} \right]$$
  
=  $-\delta(\mathbf{x}) - 3 \partial_{\mathbf{n}\mathbf{n}} \left( \frac{1}{4\pi |\mathbf{x}|} \right), \quad \mathbf{x} \in \mathbb{R}^3,$  (1.4)

with **n** =  $(n_1, n_2, n_3)^T$ , a given unit vector i.e.  $|\mathbf{n}(t)| = \sqrt{n_1^2 + n_2^2 + n_3^2} = 1$ , representing the dipole axis (or dipole moment),  $\partial_{\mathbf{n}} = \mathbf{n} \cdot \nabla$  and  $\partial_{\mathbf{nn}} = \partial_{\mathbf{n}}(\partial_{\mathbf{n}})$ . While in 2D, it is defined as [27,31]

$$U_{\rm dip}(\mathbf{x}) = -\frac{3}{2} \left( \partial_{\mathbf{n}_{\perp} \mathbf{n}_{\perp}} - n_3^2 \nabla_{\perp}^2 \right) \left( \frac{1}{2\pi |\mathbf{x}|} \right), \quad \mathbf{x} \in \mathbb{R}^2, \tag{1.5}$$

where  $\nabla_{\perp} = (\partial_x, \partial_y)^T$ ,  $\mathbf{n}_{\perp} = (n_1, n_2)^T$ ,  $\partial_{\mathbf{n}_{\perp}} = \mathbf{n}_{\perp} \cdot \nabla_{\perp}$  and  $\partial_{\mathbf{n}_{\perp}\mathbf{n}_{\perp}} = \partial_{\mathbf{n}_{\perp}}(\partial_{\mathbf{n}_{\perp}})$ . In fact, for smooth densities, the DDI potential can be reformulated via the Coulomb potential whose convolution kernel is  $U_{\text{cou}}(\mathbf{x}) = \frac{1}{2^{d-1}|\mathbf{x}|}$ . To be precise, the 3D DDI potential (1.4) is reformulated as follows

$$\begin{split} \Phi_{j}(\mathbf{x}) &= -\rho_{j} - 3 \ \partial_{\mathbf{n}} \partial_{\mathbf{n}} \left( \frac{1}{4\pi |\mathbf{x}|} * \rho_{j} \right) \\ &= -\rho_{j} - 3 \ \frac{1}{4\pi |\mathbf{x}|} * (\partial_{\mathbf{n}} \partial_{\mathbf{n}} \rho_{j}), \quad \mathbf{x} \in \mathbb{R}^{3}, \end{split}$$
(1.6)

while the 2D DDI (1.5) is rewritten as

$$\Phi_{j}(\mathbf{x}) = -\frac{3}{2} \frac{1}{2\pi |\mathbf{x}|} * \left[ \left( \partial_{\mathbf{n}_{\perp} \mathbf{n}_{\perp}} - n_{3}^{2} \nabla_{\perp}^{2} \right) \rho_{j} \right], \quad \mathbf{x} \in \mathbb{R}^{2}.$$
(1.7)

The time dependent CGPE in (1.1)-(1.2) conserve two important quantities: the total mass (or normalization) of the wave function

$$\mathcal{N}(t) := \|\Psi(\mathbf{x}, t)\|^2 = \mathcal{N}_1(t) + \mathcal{N}_2(t) \equiv \|\Psi(\mathbf{x}, 0)\|^2 = 1, \quad (1.8)$$

where  $N_j(t)$  is the mass of the *j*th component at time  $t \ge 0$ , which reads as

$$\mathcal{N}_{j}(t) := \int_{\mathbb{R}^{d}} |\psi_{j}(\mathbf{x}, t)|^{2} d\mathbf{x} \equiv \mathcal{N}_{j}(0), \quad j = 1, 2, t \ge 0,$$
(1.9)

and the energy per particle

$$\begin{split} \mathcal{E}(\Psi(\cdot, t)) &= \int_{\mathbb{R}^d} \left[ \sum_{j=1}^2 \left( \frac{1}{2} |\nabla \psi_j|^2 + V_j(\mathbf{x})|\psi_j|^2 + \frac{\beta_{jj}}{2} |\psi_j|^4 \right. \\ &+ \frac{\lambda_{ij}}{2} \phi_j |\psi_j|^2 - \Omega \psi_j^* L_z \psi_j \right) \\ &+ \frac{1}{2} (\beta_{12} + \beta_{21}) |\psi_1|^2 |\psi_2|^2 \\ &+ \frac{1}{4} (\lambda_{12} + \lambda_{21}) (\phi_1 |\psi_2|^2 + \phi_2 |\psi_1|^2) \right] d\mathbf{x} \\ &\equiv \mathcal{E}(\Psi(\cdot, 0)), \qquad t \ge 0. \end{split}$$
(1.10)

There have been extensive mathematical and numerical studies on the single-component dipolar BEC; we refer the reader to [28,30-37]. For the rotating two-component BEC without DDI, dynamics and stationary states have been studied in [38,39]and [40-42], respectively. Recently, there is growing interest from physicists for studying the properties of (non)-rotating twocomponent BEC with DDI [20-26,29,43]. However, up to now, there are few numerical studies on the rotating two-component BEC with DDI based on the CGPE (1.1)-(1.2). In this paper, we give an exhaustive mathematical study of new efficient numerical methods of the rotating two-component dipolar BECs.

To compute the dynamics, the main difficulties lie in the nonlocal DDI evaluation and proper treatment of the rotation term. As is shown before, the DDI can be computed via the Coulomb potential. On bounded rectangular domains with Dirichlet boundary conditions, the Discrete Sine Transform (DST) method applies directly [28,30]. However, the DST method requires a quite large computation domain in order to achieve a satisfactory accuracy. In 2014, Jiang et al. [44] proposed a NonUniform Fast Fourier Transform (NUFFT) solver by adopting the polar/spherical coordinates in the Fourier domain, we refer to [32,45] for extensions and applications in the context of Nonlinear Schrödinger equation (NLSE). Recently, using an accurate Gaussian-summation approximation of the convolution kernel, Zhang et al. [51] introduced an even more efficient and accurate method, which we shall refer to as GauSum solver hereafter. Both NUFFT and GauSum solver are fast algorithms with a complexity of  $O(N \log N)$  where N is the total number of grid points. Compared with the NUFFT solver, the GauSum solver is 3-5 times faster, thus it is the state-of-the-art method for applications [46]. For the rotation term, Bao et al. [47] developed a rotating Lagrangian coordinates transformation method to reformulate the rotating term into a time-dependent trapping potential, and this method allows for the implementation of high order time marching schemes [48-50]. Note that in Eulerian coordinates, additional efforts have to be made for the rotational term. In the literature, to discretize the rotational term, popular approaches either introduce ADI technique or use polar/spherical coordinate. The former method introduces extra splitting error and is complicated to be extended to higher order time marching schemes, while the later imposes artificial boundary at the origin point r = 0 and the radial direction is discretized by lower order schemes, for example finite difference method (FDM) and finite element method (FEM).

The main objectives of this paper are threefold.

1. Using the rotating Lagrangian coordinates transform [47], we reformulate the original CGPE into new equations without rotating term. Then we develop a robust and efficient numerical method to compute dynamics of the new equations by incorporating the GauSum solver [51], which is designed to compute the nonlocal DDI, into an adapted version of the time-splitting Fourier pseudospectral method. Detailed numerical results are presented that confirm the spectral accuracy in space and second order temporal accuracy of the proposed method in 2D and 3D respectively.

- 2. Develop the dynamical laws for the mass and energy, the angular momentum expectation and center of mass, together with some proofs. An analytical solution with special initial data is also presented.
- 3. Apply our method to study the dynamics of center of mass, quantized vortex lattices and non-rotating dipolar BECs in different setups. In particular, phase separation and collapse dynamics are observed numerically for the 3D cases.

The rest of the paper is organized as follows: In Section 2, we present a brief review of the Gaussian-sum method. In Section 3, we derive some dynamical laws for some physical quantities that are usually considered for the standard GPE. We then propose an efficient and robust time splitting Fourier pseudospectral numerical method for simulating the dynamics. Detailed convergence tests are presented in Section 4 to confirm the spatial and temporal convergence of our method, and some interesting numerical results are also reported. Finally, conclusions are drawn in Section 5.

#### 2. The DDI evaluation by Gaussian-sum method

Due to the confining potential  $V(\mathbf{x})$  in the CGPE (1.1)–(1.2), the density  $\rho(\mathbf{x}) := |\psi(\mathbf{x})|^2$  is usually smooth and decays exponentially fast. As is shown by (1.6) and (1.7), the DDI computation boils down to Coulomb potential evaluation. Therefore, in this section, we shall only give a brief self-contained review of the GauSum method [51] for Coulomb potential. All subscripts in the section are omitted for brevity.

For numerical calculations we have to truncate the whole space to a bounded domain with boundary conditions(BC), e.g. a square box  $\mathbf{B}_L := [-L, L]^d$  with homogeneous Dirichlet BC, then we rescale to a unit box  $\mathbf{B}_1$  by setting the density  $\rho$  to be zero outside  $\mathbf{B}_1$ . Using a smooth approximation of  $U_{\text{cou}}$  (see  $U_{\text{GS}}$  in (2.5)), the Coulomb potential is split into two integrals, i.e. the *long-range regular integral*  $I_1(\mathbf{x})$  and the *short-range singular integral*  $I_2(\mathbf{x})$ . To be precise,

$$\begin{split} \Phi(\mathbf{x}) &\approx \int_{\mathbf{B}_1} U_{\text{cou}}(\mathbf{x} - \mathbf{y}) \ \rho(\mathbf{y}) d\mathbf{y} = \int_{\mathbf{B}_2} U_{\text{cou}}(\mathbf{y}) \ \rho(\mathbf{x} - \mathbf{y}) d\mathbf{y} \qquad (2.1) \\ &= \int_{\mathbf{B}_2} U_{\text{GS}}(\mathbf{y}) \ \rho(\mathbf{x} - \mathbf{y}) d\mathbf{y} \end{split}$$

+ 
$$\int_{\mathcal{B}_{\delta}} (U_{\text{cou}}(\mathbf{y}) - U_{\text{GS}}(\mathbf{y})) \rho(\mathbf{x} - \mathbf{y}) d\mathbf{y} + I_{\delta}$$
 (2.2)

$$:= I_1(\mathbf{x}) + I_2(\mathbf{x}) + I_\delta, \qquad \mathbf{x} \in \mathbf{B}_1.$$
(2.3)

The remainder integral  $I_{\delta}$  is given explicitly as

$$I_{\delta} = \int_{\mathbf{B}_{2} \setminus \mathcal{B}_{\delta}} \left( U_{\text{cou}}(\mathbf{y}) - U_{GS}(\mathbf{y}) \right) \rho(\mathbf{x} - \mathbf{y}) d\mathbf{y}, \tag{2.4}$$

where  $\mathcal{B}_{\delta} := \{ \mathbf{x} | |\mathbf{x}| \le \delta \}$  is a very small ball centered at the origin and

$$U_{\rm GS}(\mathbf{y}) = U_{\rm GS}(|\mathbf{y}|) := \sum_{q=0}^{Q} w_q \, e^{-\tau_q^2 |\mathbf{y}|^2}, \quad Q \in \mathbb{N}^+, \tag{2.5}$$

where  $w_q$ ,  $\tau_q$  are weights and nodes. Here,  $U_{GS}$  is a very accurate approximation of  $U_{cou}$  in the interval [ $\delta$ , 2], i.e.

$$\|U_{\rm cou}(r) - U_{\rm GS}(r)\|_{L^{\infty}([\delta, 2])} \le \varepsilon_0, \tag{2.6}$$

where  $\varepsilon_0$  is the approximation error in  $L^{\infty}$  norm. For the remainder integral, we have  $|I_{\delta}| \leq C \varepsilon_0 \,\delta^d \, \|\rho\|_{L^{\infty}}$ , thus  $I_{\delta}$  is negligible for small  $\varepsilon_0$  and  $\delta$ . Therefore, it suffices to approximate the Coulomb potential by  $\Phi(\mathbf{x}) \approx I_1(\mathbf{x}) + I_2(\mathbf{x})$ . We refer to [51] for more details. To compute the *regular integral*  $I_1$ , plugging  $U_{GS}$  (2.5), we have

$$I_1(\mathbf{x}) = \sum_{q=0}^{Q} w_q \int_{\mathbf{B}_2} e^{-\tau_q^2 |\mathbf{y}|^2} \rho(\mathbf{x} - \mathbf{y}) d\mathbf{y}, \qquad \mathbf{x} \in \mathbf{B}_1.$$
(2.7)

The density  $\rho(\mathbf{x} - \mathbf{y}), \mathbf{x} \in \mathbf{B}_1, \mathbf{y} \in \mathbf{B}_2$  is well approximated by Fourier series as follows

$$\rho(\mathbf{z}) \approx \sum_{\mathbf{k}} \widehat{\rho}_{\mathbf{k}} \prod_{j=1}^{d} e^{\frac{2\pi i k_j}{6} (z^{(j)} + 3)}, \qquad \mathbf{z} = (z^{(1)}, \dots, z^{(d)}) \in \mathbf{B}_3.$$
(2.8)

For the Fourier coefficients  $\hat{\rho}_{\mathbf{k}}$ , we have

$$\widehat{\rho}_{\mathbf{k}} = \frac{1}{|\mathbf{B}_3|} \int_{\mathbf{B}_3} \rho(\mathbf{z}) \prod_{j=1}^d e^{\frac{-2\pi i \ k_j}{6} (z^{(j)} + 3)} d\mathbf{z},$$
(2.9)

where  $|\mathbf{B}_3| = \prod_{j=1}^d (b_j - a_j)$  is the volume. Careful calculations leads to

$$I_1(\mathbf{x}) = \sum_{\mathbf{k}} \widehat{\rho}_{\mathbf{k}} \left( \sum_{q=0}^{Q} w_q G_{\mathbf{k}}^q \right) \prod_{j=1}^d e^{\frac{2\pi i k_j}{b_j - a_j} (\mathbf{x}^{(j)} - a_j)},$$
(2.10)

where

$$G_{\mathbf{k}}^{q} = \prod_{j=1}^{d} \int_{-2}^{2} e^{-\tau_{q}^{2} |y^{(j)}|^{2}} e^{\frac{-2\pi i k_{j} y^{(j)}}{b_{j} - a_{j}}} dy^{(j)}, \qquad (2.11)$$

can be pre-computed once for all if the computation grid remains unchanged.

For the *near-field correction integral*  $I_2$ , the density function  $\rho_{\mathbf{x}}(\mathbf{y}) := \rho(\mathbf{x} - \mathbf{y})$  is approximated by a low-order Taylor expansion within  $\mathcal{B}_{\delta}$  as follows

$$\rho_{\mathbf{x}}(\mathbf{y}) \approx P_{\mathbf{x}}(\mathbf{y}) = \rho_{\mathbf{x}}(\mathbf{0}) + \sum_{j=1}^{d} \frac{\partial \rho_{\mathbf{x}}(\mathbf{0})}{\partial y_{j}} y_{j}$$
$$+ \frac{1}{2} \sum_{j,k=1}^{d} \frac{\partial^{2} \rho_{\mathbf{x}}(\mathbf{0})}{\partial y_{j} \partial y_{k}} y_{j} y_{k} + \frac{1}{6} \sum_{j,k,\ell=1}^{d} \frac{\partial^{3} \rho_{\mathbf{x}}(\mathbf{0})}{\partial y_{j} \partial y_{k} \partial y_{\ell}} y_{j} y_{k} y_{\ell}.$$
(2.12)

We then integrate it in spherical/polar coordinates. The computation boils down to a multiplication of the Laplacian  $\Delta \rho$  since the contributions of the odd derivatives in (2.12) and off-diagonal components of the Hessian vanish. Derivatives of  $\rho$  are computed via its Fourier series. For more details, we refer the reader to [46,51].

The GauSum method achieves a spectral accuracy and is as efficient as FFT algorithms within  $O(N \log N)$  arithmetic operations [51]. The algorithm has been implemented for DDI [51] and applied in the studies of fractional Schrödinger equations [46].

#### 3. Properties of the dynamics and the numerical method

In this section, we first present conservation laws for some commonly used quantities analogous to the rotating CGPE without DDI. Then, we extend the rotating Lagrangian coordinate transform proposed for the GPE without DDI in [47]. In the rotating Lagrangian coordinates, the rotation term vanishes, instead the potential becomes time-dependent. For the new equation, we shall propose a time-splitting Fourier spectral method incorporated with the GauSum solver to compute the dynamics.

#### 3.1. Properties of the dynamics

Here we study the dynamical properties of the mass, energy, angular momentum expectation and center of mass. The dynamical laws can be used as benchmarks to test the numerical methods and are briefly listed here. For details, one can prove in an analogous way to the one component [47,52] or two-component without DDI [30].

**Mass and energy.** The CGPE (1.1)-(1.2) conserves the mass (1.9) and energy (1.10), i.e.

$$\mathcal{N}_{1}(t) = \mathcal{N}_{1}(t=0), \quad \mathcal{N}_{2}(t) = \mathcal{N}_{2}(t=0),$$
  
$$\mathcal{N}(t) := (\mathcal{N}_{1} + \mathcal{N}_{2})(t) = \mathcal{N}(t=0), \quad \mathcal{E}(t) = \mathcal{E}(t=0).$$
(3.1)

**Angular momentum expectation.** The *angular momentum expectation* for each component and the total angular momentum are defined respectively as

$$\langle L_z \rangle_j(t) = \int_{\mathbb{R}^d} \psi_j^*(\mathbf{x}, t) L_z \psi_j(\mathbf{x}, t) \, d\mathbf{x}, \quad j = 1, 2, \langle L_z \rangle(t) = \langle L_z \rangle_1(t) + \langle L_z \rangle_2(t), \quad t \ge 0.$$
 (3.2)

**Lemma 3.1.** If  $V_j(\mathbf{x})$  reads as the harmonic potential, we have for j = 1, 2 and  $k_j = 3 - j$ 

$$\frac{d}{dt} \langle L_z \rangle_j(t) = w_j^- \int_{\mathbb{R}^d} xy |\psi_j|^2 d\mathbf{x} 
+ \int_{\mathbb{R}^d} |\psi_j|^2 (y\partial_x - x\partial_y) \Big(\beta_{jk_j} |\psi_{k_j}|^2 + \sum_{k=1}^2 \lambda_{jk} \Phi_k(\mathbf{x}, t) \Big) d\mathbf{x}.$$
(3.3)

Moreover, if additionally  $\beta_{12} = \beta_{21}$ , we have

$$\frac{d}{dt} \langle L_z \rangle(t) = \sum_{j=1}^2 w_j^- \int_{\mathbb{R}^d} xy |\psi_j|^2 d\mathbf{x} + \sum_{j,k=1}^2 \lambda_{jk} \int_{\mathbb{R}^d} |\psi_j|^2 (y\partial_x - x\partial_y) \Phi_k(\mathbf{x}, t) d\mathbf{x}.$$
(3.4)

Here  $w_j^- = \gamma_{x,j}^2 - \gamma_{y,j}^2$ . This implies that the total angular momentum expectation  $\langle L_z \rangle(t)$  is conserved, i.e.

$$\langle L_z \rangle(t) = \langle L_z \rangle(0), \qquad t \ge 0,$$
 (3.5)

if  $\gamma_{x,j} = \gamma_{y,j}$  and one of the following condition holds: (i)  $\lambda_{11} = \lambda_{12} = \lambda_{21} = \lambda_{22} = 0$ . (ii)  $\lambda_{12} = \lambda_{21}$  and the dipole axises parallel to the *z*-axis, i.e.  $\mathbf{n}_1 = \mathbf{n}_2 = (0, 0, 1)^T$ . Moreover, the angular momentum expectation for each component is also conserved, i.e.

$$\langle L_z \rangle_j(t) = \langle L_z \rangle_j(0), \qquad j = 1, 2, \quad t \ge 0, \tag{3.6}$$

*if additionally provided*  $\lambda_{12} = \lambda_{21} = \beta_{12} = \beta_{21} = 0$ .

**Proof.** Let us take a close look at the DDI term in (3.4). Using the Plancherel's formula, we have

$$\langle \rho_j, (y\partial_x - x\partial_y)\Phi_k \rangle := \int_{\mathbb{R}^d} |\psi_j|^2 (-\partial_\theta \Phi_k) d\mathbf{x}$$
  
=  $\frac{1}{(2\pi)^d} \langle \widehat{\rho_j}, \widehat{-\partial_\theta \Phi_k} \rangle$  (3.7)

$$= \frac{1}{(2\pi)^d} \langle \widehat{\rho}_j, -\partial_{\theta_{\xi}} \widehat{U_{\text{dip}}} \widehat{\rho}_k \rangle$$
(3.8)

where the Fourier transform is defined as  $\widehat{f}(\xi) = \int_{\mathbb{R}^d} f(\mathbf{x}) e^{-i\xi \cdot \mathbf{x}} d\mathbf{x}$ ,

$$\widehat{U}_{dip}(\xi) = \begin{cases} -1 + \frac{3(\mathbf{n} \cdot \xi)^2}{|\xi|^2}, & d = 3, \\ \frac{2[(\mathbf{n}_{\perp} \cdot \xi)^2 - n_3^2 |\xi|^2]}{2|\xi|}, & d = 2, \end{cases}$$
(3.9)

and  $\theta$ ,  $\theta_{\xi}$  are the azimuth angle in physical/Fourier space respectively. For  $\mathbf{n} = (0, 0, 1)^{T}$ , it is easy to see that  $\widehat{U}_{dip}(\xi)$  is cylindrical/polar symmetric in 3D and 2D respectively, and we have

$$\langle \rho_j, -\partial_\theta \Phi_k \rangle = \frac{1}{(2\pi)^d} \langle \widehat{\rho_j}, -\partial_{\theta_{\xi}} (\widehat{U_{\text{dip}}} \widehat{\rho_k}) \rangle$$

$$= \frac{1}{(2\pi)^d} \langle \widehat{U_{\text{dip}}} \widehat{\rho_j}, -\partial_{\theta_{\xi}} \widehat{\rho_k} \rangle$$
(3.10)

$$= \langle \Phi_j, -\partial_\theta \rho_k \rangle = \langle \partial_\theta \Phi_j, \rho_k \rangle \tag{3.11}$$

$$= -\langle \rho_k, -\partial_\theta \Phi_j \rangle. \tag{3.12}$$

The proof is then completed due to the above anti-symmetric property in the index (j, k).

Center of mass. The total center of mass is defined as

$$\mathbf{x}_{c}(t) = \int_{\mathbb{R}^{d}} \mathbf{x} \left| \Psi(\mathbf{x}, t) \right|^{2} d\mathbf{x} \eqqcolon \mathbf{x}_{c,1}(t) + \mathbf{x}_{c,2}(t), \quad t \ge 0, \quad (3.13)$$

where  $\mathbf{x}_{c,j}(t)$  is defined as the (scaled) center of mass of the *j*th component and reads as follows

$$\mathbf{x}_{c,j}(t) = \int_{\mathbb{R}^d} \mathbf{x} \, |\psi_j(\mathbf{x},t)|^2 d\mathbf{x}, \quad j = 1, \ 2, \quad t \ge 0.$$
(3.14)

**Lemma 3.2.** If  $V_j(\mathbf{x})$  reads as the harmonic potential, we have for j = 1, 2 and  $k_j = 3 - j$ 

$$\begin{split} \ddot{\mathbf{x}}_{c,j} &- 2\Omega J_d \dot{\mathbf{x}}_{c,j} + (\Lambda_{d,j} + \Omega^2 J_d^2) \mathbf{x}_{c,j} \\ &= \int_{\mathbb{R}^d} \left( \beta_{j,k_j} |\psi_{k_j}|^2 + \lambda_{j,k_j} \Phi_{k_j} \right) \nabla |\psi_j|^2 \, d\mathbf{x}, \end{split}$$
(3.15)  
$$\mathbf{x}_{c,j}^0 &= \int_{-\infty}^{\infty} \mathbf{x} |\psi_j^0(\mathbf{x})|^2 d\mathbf{x}, \end{split}$$

$$\dot{\mathbf{x}}_{c,j}^{0} = \int_{\mathbb{R}^d} \operatorname{Im}(\bar{\psi}_j^0 \nabla \psi_j^0) d\mathbf{x} + \Omega J_d \mathbf{x}_{c,j}^0, \qquad (3.16)$$

where,

$$J_{d} = \begin{cases} \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}, \\ \begin{pmatrix} J_{2} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} \end{pmatrix}, \\ \Lambda_{d,j} = \begin{cases} \begin{pmatrix} \gamma_{x,j}^{2} & 0 \\ 0 & \gamma_{y,j}^{2} \end{pmatrix}, & d = 2, \\ \begin{pmatrix} \Lambda_{j,2} & \mathbf{0} \\ \mathbf{0} & \gamma_{z,j}^{2} \end{pmatrix}, & d = 3. \end{cases}$$
(3.17)

Moreover, if  $V_1(\mathbf{x}) = V_2(\mathbf{x})$ ,  $\beta_{12} = \beta_{21}$  and  $\lambda_{12} = \lambda_{21}$ , we have

$$\ddot{\mathbf{x}}_c - 2\Omega J_d \dot{\mathbf{x}}_c + (\Lambda_{d,1} + \Omega^2 J_d^2) \mathbf{x}_c = \mathbf{0}, \qquad (3.18)$$

$$\mathbf{x}_{c}^{0} = \mathbf{x}_{c,1}^{0} + \mathbf{x}_{c,2}^{0}, \qquad \dot{\mathbf{x}}_{c}^{0} = \dot{\mathbf{x}}_{c,1}^{0} + \dot{\mathbf{x}}_{c,2}^{0}.$$
(3.19)

An analytical solution for special initial data. An interesting application of the dynamic law (3.18) for the total center of mass is that under some circumstances we can construct an analytical solution to the CGPE. Precisely speaking, suppose the initial condition  $\psi_i^0$  in (1.2) is chosen as

$$\psi_j^0(\mathbf{x}) = \phi_j^s(\mathbf{x} - \mathbf{x}_0), \qquad \mathbf{x} \in \mathbb{R}^d, \tag{3.20}$$

where  $\mathbf{x}_0 \in \mathbb{R}^d$  is a given point and  $\phi_j^s$  (j = 1, 2) is a stationary state of the CGPE, i.e.

$$\mu_{j}^{s}\phi_{j}^{s} = \left[-\frac{1}{2}\nabla^{2} + V_{j}(\mathbf{x}) - \Omega L_{z} + \sum_{k=1}^{2} \left(\beta_{jk}|\phi_{k}^{s}|^{2} + \lambda_{jk}U_{dip} * |\phi_{k}^{s}|^{2}\right)\right]\phi_{j}^{s}, \qquad (3.21)$$

$$\int_{\mathbb{R}^d} |\phi_j^s|^2 \, d\mathbf{x} = 1, \ \ j = 1, 2$$
(3.22)

where  $\mu_j^s \in \mathbb{R}$  (j = 1, 2) are the chemical potentials. With this initial value, and suppose  $V_1(\mathbf{x}) = V_2(\mathbf{x})$ , the exact solution of the CGPE with harmonic potential can be constructed as

$$\psi_j(\mathbf{x},t) = \phi_j^{\mathrm{s}}(\mathbf{x} - \mathbf{x}_c(t))e^{-i\mu_j^{\mathrm{s}}t}e^{iw(\mathbf{x},t)}, \quad \mathbf{x} \in \mathbb{R}^d, \quad t \ge 0,$$
(3.23)

where  $w(\mathbf{x}, t)$  is linear in  $\mathbf{x}$ , i.e.

$$w(\mathbf{x},t) = \mathbf{c}_1(t) \cdot \mathbf{x} + \mathbf{c}_2(t), \quad \mathbf{x} \in \mathbb{R}^d, \quad t \ge 0,$$
(3.24)

with some functions  $\mathbf{c}_1(t)$ ,  $\mathbf{c}_2(t)$ , and  $\mathbf{x}(t)$  satisfying the ODE (3.18) with initial condition

$$\mathbf{x}_c^0 = \mathbf{x}^0, \qquad \dot{\mathbf{x}}_c^0 = -\Omega J_d \mathbf{x}^0.$$
(3.25)

## 3.2. Numerical method

## 3.2.1. CGPE under rotating Lagrangian coordinates

In this section, we first introduce a rotating Lagrangian coordinate and then reformulate the CGPE (1.1)-(1.2) in the new coordinate system. For any time  $t \ge 0$ , let  $\mathbf{A}_d(t)$  be an orthogonal rotational matrix in  $\mathbb{R}^d$  defined as [47,50]

$$\mathbf{A}_{d}(t) = \begin{pmatrix} \cos(\Omega t) & \sin(\Omega t) \\ -\sin(\Omega t) & \cos(\Omega t) \end{pmatrix}, \quad \text{if } d = 2,$$
$$\mathbf{A}_{d}(t) = \begin{pmatrix} \mathbf{A}_{2}(t) & \mathbf{0} \\ \mathbf{0} & 1 \end{pmatrix}, \quad \text{if } d = 3.$$
(3.26)

It is easy to verify that  $\mathbf{A}_d^{-1}(t) = \mathbf{A}_d^T(t)$  for any  $t \ge 0$  and A(0) = I with I the identity matrix. For  $\forall t \ge 0$ , the *rotating Lagrangian* coordinates  $\widetilde{\mathbf{x}}$  is defined as

$$\widetilde{\mathbf{x}} = \mathbf{A}_d^{-1}(t)\mathbf{x} = \mathbf{A}_d^T(t)\mathbf{x} \quad \Longleftrightarrow \quad \mathbf{x} = \mathbf{A}_d(t)\widetilde{\mathbf{x}}, \mathbf{x} \in \mathbb{R}^d, \quad t \ge 0.$$
(3.27)

Denote the wave function in the new coordinates as  $\phi_i(\tilde{\mathbf{x}}, t)$ :

$$\begin{split} \phi_{j}(\widetilde{\mathbf{x}}, t) &:= \psi_{j}(\mathbf{x}, t) = \psi_{j}\left(\mathbf{A}_{d}(t)\widetilde{\mathbf{x}}, t\right), \quad j = 1, 2\\ \widetilde{\mathbf{x}} \in \mathbb{R}^{d}, \quad t \geq 0. \end{split}$$
(3.28)

By simple calculation, we have

$$\begin{split} i\partial_t \phi_j(\widetilde{\mathbf{x}}, t) &= i\partial_t \psi_j(\mathbf{x}, t) + i\nabla_{\mathbf{x}} \psi_j(\mathbf{x}, t) \cdot \left(\dot{\mathbf{A}}_d(t)\widetilde{\mathbf{x}}\right) \\ &= i\partial_t \psi_j(\mathbf{x}, t) + \Omega L_z \psi_j(\mathbf{x}, t), \\ \nabla_{\widetilde{\mathbf{x}}} \phi_j(\widetilde{\mathbf{x}}, t) &= A_d^{-1}(t) \nabla_{\mathbf{x}} \psi_j(\mathbf{x}, t), \quad \nabla_{\widetilde{\mathbf{x}}}^2 \phi_j(\widetilde{\mathbf{x}}, t) = \nabla_{\mathbf{x}}^2 \psi_j(\mathbf{x}, t), \\ &\mathbf{x} \in \mathbb{R}^d, \quad t \ge 0. \end{split}$$

Substituting the above derivatives into (1.1)–(1.2) leads to the following *d*-dimensional CGPE in the rotating Lagrangian coordinates  $\tilde{\mathbf{x}}$ , for j = 1, 2

$$i\frac{\partial\phi_{j}(\widetilde{\mathbf{x}},t)}{\partial t} = \left[-\frac{1}{2}\nabla^{2} + \mathcal{W}_{j}(\widetilde{\mathbf{x}},t) + \sum_{k=1}^{2} (\beta_{jk}|\phi_{k}|^{2} + \lambda_{jk}\widetilde{\Phi}_{k})\right]\phi_{j},$$
  

$$\widetilde{\mathbf{x}} \in \mathbb{R}^{d}, \quad t > 0, \qquad (3.29)$$
  

$$\widetilde{\Phi}_{k}(\widetilde{\mathbf{x}},t) = \widetilde{U}_{dip} * |\phi_{k}|^{2}, \qquad \phi_{j}(\widetilde{\mathbf{x}},0) \coloneqq \phi_{j}^{0}(\widetilde{\mathbf{x}}) = \psi_{j}^{0}(\mathbf{x}),$$
  

$$\widetilde{\mathbf{x}} = \mathbf{x} \in \mathbb{R}^{d}. \qquad (3.30)$$

Here,  $W_j(\widetilde{\mathbf{x}}, t) = V_j(\mathbf{A}_d(t)\widetilde{\mathbf{x}})$  (j = 1, 2) and the DDI kernel  $\widetilde{U}_{dip}(\widetilde{\mathbf{x}}, t)$  reads as

$$\widetilde{U}_{dip}(\widetilde{\mathbf{x}}, t) = \begin{cases} -\delta(\widetilde{\mathbf{x}}) - 3 \,\partial_{\mathbf{m}(t)\mathbf{m}(t)} \left(\frac{1}{4\pi \,|\widetilde{\mathbf{x}}|}\right), \\ d = 3, \\ -\frac{3}{2} \left(\partial_{\mathbf{m}_{\perp}(t)\mathbf{m}_{\perp}(t)} - m_3^2 \,\nabla_{\perp}^2\right) \left(\frac{1}{2\pi \,|\widetilde{\mathbf{x}}|}\right), \\ d = 2, \end{cases}$$
(3.31)

with  $\mathbf{m}(t) \in \mathbb{R}^3$  defined as  $\mathbf{m}(t) = \mathbf{A}_d^{-1}(t)\mathbf{n} = (m_1(t), m_2(t), m_3(t))^T$  and  $\mathbf{m}_{\perp}(t) := (m_1(t), m_2(t))^T$ .

In rotating Lagrangian coordinates, the energy associated with the CGPE (3.29)–(3.30) is defined as

$$\widetilde{\varepsilon}(t) = \sum_{j=1}^{2} \int_{\mathbb{R}^{d}} \left[ \frac{1}{2} |\nabla \phi_{j}|^{2} + W_{j}(\widetilde{\mathbf{x}}, t)|\phi_{j}|^{2} + \sum_{k=1}^{2} \left( \frac{\beta_{jk}}{2} |\phi_{k}|^{2} + \frac{\lambda_{jk}}{2} \phi_{k} \right) |\phi_{j}|^{2} \right] d\widetilde{\mathbf{x}} - \sum_{j=1}^{2} \int_{\mathbb{R}^{d}} \int_{0}^{t} \left[ \partial_{\tau} W_{j}(\widetilde{\mathbf{x}}, \tau) d\tau + \sum_{k=1}^{2} \frac{\lambda_{jk}}{2} (\partial_{\tau} \widetilde{U}_{dip}) * |\phi_{k}|^{2} \right] |\phi_{j}|^{2} d\widetilde{\mathbf{x}} =: \widetilde{\varepsilon}_{kin}(t) + \widetilde{\varepsilon}_{pot}(t) + \widetilde{\varepsilon}_{short}(t) + \widetilde{\varepsilon}_{dip}(t) + \widetilde{\varepsilon}_{extra}(t), \quad (3.32)$$

where

$$\begin{split} \widetilde{\varepsilon}_{kin}(t) &= \frac{1}{2} \int_{\mathbb{R}^d} \left[ |\nabla \phi_1|^2 + |\nabla \phi_2|^2 \right] d\widetilde{\mathbf{x}}, \\ \widetilde{\varepsilon}_{pot}(t) &= \int_{\mathbb{R}^d} \left[ W_1(\widetilde{\mathbf{x}}, t) |\phi_1|^2 + W_2(\widetilde{\mathbf{x}}, t) |\phi_2|^2 \right] d\widetilde{\mathbf{x}}, \\ \widetilde{\varepsilon}_{short}(t) &= \frac{1}{2} \sum_{j,k=1}^2 \beta_{jk} \int_{\mathbb{R}^d} |\phi_j|^2 |\phi_k|^2 d\widetilde{\mathbf{x}}, \\ \widetilde{\varepsilon}_{dip}(t) &= \frac{1}{2} \sum_{j,k=1}^2 \lambda_{jk} \int_{\mathbb{R}^d} \Phi_k |\phi_j|^2 d\widetilde{\mathbf{x}}, \\ \widetilde{\varepsilon}_{extra}(t) &= -\sum_{j=1}^2 \int_{\mathbb{R}^d} \int_0^t \left[ \partial_\tau W_j(\widetilde{\mathbf{x}}, \tau) d\tau \right] \\ &+ \sum_{k=1}^2 \frac{\lambda_{jk}}{2} (\partial_\tau \widetilde{U}_{dip}) * |\phi_k|^2 \right] |\phi_j|^2 d\widetilde{\mathbf{x}}, \end{split}$$

and

$$\partial_{t} \widetilde{U}_{dip}(\widetilde{\mathbf{x}}, t) = -3 \begin{cases} 2 \,\partial_{\widetilde{\mathbf{m}}(t)\mathbf{m}(t)} \left(\frac{1}{4\pi \,|\widetilde{\mathbf{x}}|}\right), & d = 3, \\ \\ \partial_{\widetilde{\mathbf{m}}_{\perp}(t)\mathbf{m}_{\perp}(t)} \left(\frac{1}{2\pi \,|\widetilde{\mathbf{x}}|}\right), & d = 2. \end{cases}$$
(3.33)

**Remark 3.1.** If  $V_j(\mathbf{x})$  is a harmonic potential as defined in (1.3), then  $W_j(\mathbf{\tilde{x}}, t)$  has the form

$$\mathcal{W}_{j}(\widetilde{\mathbf{x}}, t) = \frac{w_{j}^{+}}{4} (\widetilde{\mathbf{x}}^{2} + \widetilde{\mathbf{y}}^{2}) + \frac{w_{j}^{-}}{4} [(\widetilde{\mathbf{x}}^{2} - \widetilde{\mathbf{y}}^{2}) \cos(2\Omega t) + 2\widetilde{\mathbf{x}}\widetilde{\mathbf{y}} \sin(2\Omega t)] + \begin{cases} 0 & d = 2, \\ \frac{1}{2}\gamma_{z,j}^{2}\widetilde{\mathbf{z}}^{2}, & d = 3, \end{cases}$$
(3.34)

where  $w_j^+ = \gamma_{x,j}^2 + \gamma_{y,j}^2$  and  $w_j^- = \gamma_{x,j}^2 - \gamma_{y,j}^2$ . Therefore, when the external potential is either a box-potential or a harmonic potential which is radially symmetric in two dimensions (2D) or cylindrically

symmetric in three dimensions (3D), i.e.  $\gamma_{x,j} = \gamma_{y,j} := \gamma_{r,j}$ , the potential  $W_j(\widetilde{\mathbf{x}}, t)$  becomes time-independent.

Compared to (1.1)-(1.2), the rotating term now vanishes in the new CGPE (3.29)-(3.30). Instead, the trapping potential and DDI kernel now become time-dependent. The absence of rotating term now allows us to develop an efficient method to solve (3.29)-(3.30).

## 3.2.2. Time splitting Fourier pseudospectral method

Here we shall consider the new equations (3.29)-(3.30). Due to the trapping potential, the wave functions decay exponentially. Therefore it suffices to truncate the problem into a sufficiently large bounded computational domain  $\mathcal{D} = [a, b] \times [c, e] \times [f, g]$  if d = 3, or  $\mathcal{D} = [a, b] \times [c, e]$  if d = 2. From  $t = t_n$  to  $t = t_{n+1} := t_n + \Delta t$ , the CGPE will be solved in two steps, i.e. for j = 1, 2 one first solves

$$i\partial_t \phi_j(\widetilde{\mathbf{x}}, t) = -\frac{1}{2} \nabla^2 \phi_j(\widetilde{\mathbf{x}}, t), \quad \widetilde{\mathbf{x}} \in \mathcal{D}, \quad t_n \le t \le t_{n+1},$$
 (3.35)

with periodic boundary conditions on the boundary  $\partial D$  for a time step of length  $\Delta t$ , then solves

$$i\frac{\partial\phi_{j}(\widetilde{\mathbf{x}},t)}{\partial t} = \left[\mathcal{W}_{j}(\widetilde{\mathbf{x}},t) + \sum_{k=1}^{2} \left(\beta_{jk}|\phi_{k}|^{2} + \lambda_{jk}\widetilde{\Phi}_{k}\right)\right]\phi_{j}$$
$$\widetilde{\mathbf{x}} \in \mathcal{D}, \quad t_{n} \leq t \leq t_{n+1},$$
(3.36)

$$\widetilde{\boldsymbol{\Phi}}_{k}(\widetilde{\mathbf{x}},t) = \left(\widetilde{\boldsymbol{U}}_{\text{dip}} * \widetilde{\rho}_{k}\right)(\widetilde{\mathbf{x}},t),$$

$$k = 1, 2, \qquad \widetilde{\mathbf{x}} \in \mathcal{D}, \quad t_{n} \le t \le t_{n+1}, \qquad (3.37)$$

for the same time step. Here,  $\tilde{\rho}_k(\mathbf{\tilde{x}}, t) = |\phi_k(\mathbf{\tilde{x}}, t)|^2$  if  $\mathbf{\tilde{x}} \in D$ and  $\tilde{\rho}_k(\mathbf{\tilde{x}}, t) = 0$  otherwise. The linear subproblem (3.35) will be discretized in space by the Fourier pseudospectral method and integrated in time exactly in the phase space, while the nonlinear subproblem (3.36)–(3.37) preserves the density point-wise, i.e.  $|\phi_j(\mathbf{\tilde{x}}, t)|^2 \equiv |\phi_j(\mathbf{\tilde{x}}, t = t_n)|^2 = |\phi_j^n(\mathbf{\tilde{x}})|^2$ , and it can be integrated exactly as

$$\phi_{j}(\mathbf{x}, t) = \exp\left\{-i\left[P_{j}(\mathbf{x}, t) + \sum_{k=1}^{2} \left(\beta_{jk} |\phi_{k}^{n}|^{2}(t-t_{n}) + \lambda_{jk} \varphi_{k}(\widetilde{\mathbf{x}}, t)\right)\right]\right\},$$
(3.38)

$$\varphi_{k}(\widetilde{\mathbf{x}}, t) = \int_{\mathbb{R}^{d}} \widetilde{\kappa}(\widetilde{\mathbf{x}} - \widetilde{\mathbf{y}}, t) \rho_{k}(\widetilde{\mathbf{y}}, t_{n}) d\widetilde{\mathbf{y}},$$
  
$$\widetilde{\mathbf{x}} \in \mathcal{D}, \quad t_{n} \le t \le t_{n+1},$$
(3.39)

where the time-dependent kernel  $\widetilde{\mathcal{K}}(\mathbf{\tilde{x}}, t)$  has the form

$$\widetilde{\mathcal{K}}(\widetilde{\mathbf{x}},t) = \int_{t_n}^{t} \widetilde{U}_{dip}(\widetilde{\mathbf{x}},\tau) d\tau$$

$$= \begin{cases} -\delta(\widetilde{\mathbf{x}})(t-t_n) - 3\widetilde{L}_3(t)(\frac{1}{4\pi |\widetilde{\mathbf{x}}|}), & \text{3D DDI,} \\ -\frac{3}{2}\widetilde{L}_2(t)(\frac{1}{2\pi |\widetilde{\mathbf{x}}|}), & \text{2D DDI.} \end{cases}$$
(3.40)

Here, the differential operators  $\widetilde{L}_3(t) = \int_{t_n}^t \partial_{\mathbf{m}(\tau)\mathbf{m}(\tau)} d\tau$  and  $\widetilde{L}_2(t) = \int_{t_n}^t \partial_{\mathbf{m}_{\perp}(\tau)\mathbf{m}_{\perp}(\tau)} - m_3^2 \nabla_{\perp}^2 d\tau$  actually can be integrated analytically and have explicit expressions, one can refer to Section 4.1 in [47] for details. The GauSum solver is then applied to evaluate the nonlocal potential  $\varphi(\mathbf{\widetilde{x}}, t)$  (3.39). In addition,

$$P_{j}(\widetilde{\mathbf{x}}, t) = \int_{t_{n}}^{t} \mathcal{W}_{j}(\widetilde{\mathbf{x}}, \tau) d\tau.$$
(3.41)

**Remark 3.2.** If  $V_j(\mathbf{x})$  (j = 1, 2) is a harmonic potential as defined in (1.3), i.e.  $W_j$  reads as (3.34), then the integral in (3.41) can be evaluated analytically, i.e.

$$\int_{t_n}^t \mathcal{W}_j(\widetilde{\mathbf{x}}, \tau) d\tau = \frac{w_j^+(\widetilde{\mathbf{x}}^2 + \widetilde{y}^2)}{4} (t - t_n) + \frac{w_j^-}{8\Omega} \Big[ (\widetilde{\mathbf{x}}^2 - \widetilde{y}^2) \big( \sin(2\Omega t) - \sin(2\Omega t_n) \big) \quad (3.42) - 2\widetilde{\mathbf{x}} \widetilde{\mathbf{y}} \big( \cos(2\Omega t) - \cos(2\Omega t_n) \big) \Big] + \begin{cases} 0, & d = 2, \\ \frac{1}{2} \gamma_{z,j}^2 \widetilde{z}^2 (t - t_n), & d = 3. \end{cases} \quad (3.43)$$

For a general potential  $V_j(\mathbf{x})$ , if the integral in (3.41) cannot be found analytically, numerical quadratures such as Trapezoidal rule or Simpson's rule can be applied [47,53].

To simplify the presentation, we will only present the scheme for the 3D case. As for the 2D case, one can modify the algorithm straightforward. Let *L*, *M*, *N* be even positive integers, choose  $h_{\tilde{x}} = \frac{b-a}{L}$ ,  $h_{\tilde{y}} = \frac{d-c}{M}$  and  $h_{\tilde{z}} = \frac{f-e}{N}$  as the spatial mesh sizes in  $\tilde{x}$ -,  $\tilde{y}$ -, and  $\tilde{z}$ - directions, respectively. Define the index and grid points sets as

$$\begin{split} \mathcal{T}_{LMN} &= \{(l,k,m) \mid 0 \le l \le L, \ 0 \le k \le M, \ 0 \le m \le N\}, \\ \widetilde{\mathcal{T}}_{LMN} &= \left\{ (p,q,r) \mid -\frac{L}{2} \le p \le \frac{L}{2} - 1, \ -\frac{M}{2} \le q \le \frac{M}{2} - 1, \\ &- \frac{N}{2} \le r \le \frac{N}{2} - 1 \right\}, \\ \mathcal{G}_{\widetilde{x}\widetilde{y}\widetilde{z}} &= \{(\widetilde{x}_{l},\widetilde{y}_{k},\widetilde{z}_{m}) \\ &=: \ (lh_{\widetilde{x}} + a, \ kh_{\widetilde{y}} + c, \ mh_{\widetilde{z}} + e), \ (l,k,m) \in \mathcal{T}_{LMN} \right\}. \end{split}$$

Define the functions

$$W_{pqr}(\widetilde{x},\widetilde{y},\widetilde{z}) = e^{i\mu_p^{\widetilde{\chi}}(\widetilde{x}-a)} e^{i\mu_q^{\gamma}(\widetilde{y}-c)} e^{i\mu_r^{\widetilde{\chi}}(\widetilde{z}-e)}, \quad (p,q,r) \in \widetilde{\mathcal{T}}_{LMN},$$
with

$$\mu_p^{\widetilde{\chi}} = rac{2\pi p}{b-a}, \ \ \mu_q^{\widetilde{y}} = rac{2\pi q}{d-c}, \ \ \mu_r^{\widetilde{z}} = rac{2\pi r}{f-e}, \ \ \ (p,q,r) \in \widetilde{\mathcal{T}}_{LMN}.$$

Let  $f_{j,lkm}^n$   $(j = 1, 2, f_j = \phi_j, \varphi_j \text{ or } P_j)$  be the numerical approximation of  $f_j(\widetilde{x}_l, \widetilde{y}_k, \widetilde{z}_m, t_n)$  for  $(l, k, m) \in \mathcal{T}_{LMN}, n \ge 0$  and denote  $\phi_j^n$  as the solution vector at time  $t = t_n$  with components  $\{\phi_{j,lkm}^n, (l, k, m) \in \mathcal{T}_{LMN}\}$ . Taking the initial data as  $\phi_{j,lkm}^0 = \phi_j^0(\widetilde{x}_l, \widetilde{y}_k, \widetilde{z}_m)$  for  $(l, k, m) \in \mathcal{T}_{LMN}$ , a second-order Time Splitting Fourier Pseudospectral (TSFP) method to solve the CGPE (3.29)–(3.30) reads as follows:

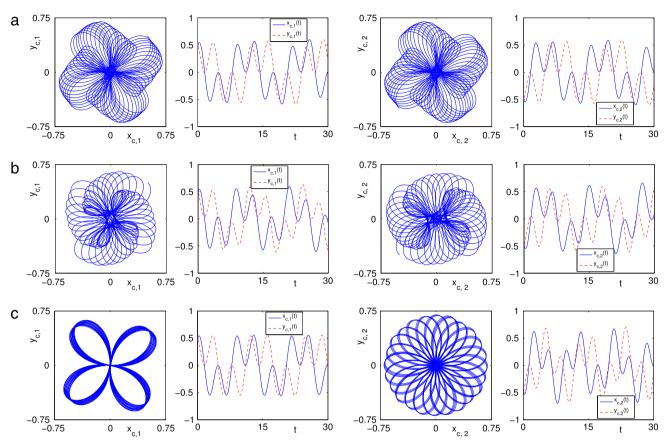
$$\phi_{j,lkm}^{(1)} = \sum_{p=-L/2}^{L/2-1} \sum_{q=-M/2}^{M/2-1} \sum_{r=-N/2}^{N/2-1} e^{-\frac{i\Delta t}{4} \left[ (\mu_p^{\widetilde{\chi}})^2 + (\mu_q^{\widetilde{\chi}})^2 + (\mu_r^{\widetilde{\chi}})^2 \right]} (\widehat{\phi_j^n})_{pqr} \times W_{pqr}(\widetilde{\chi}_l, \widetilde{y}_k, \widetilde{z}_m),$$
(3.44)  
$$\phi_{l,u}^{(2)} = \phi_{l,u}^{(1)} \exp \left\{ -i \left[ \Delta t \sum \left( \beta_{ku} | \phi_{l,u}^{(1)} | \right]^2 + \lambda_{iv} \sigma_{u,u}^{n+1} \right) \right\}$$

$$+ P_{j,lkm}^{n+1} \bigg] \bigg\},$$

$$(3.45)$$

$$\phi_{j,lkm}^{n+1} = \sum_{p=-L/2}^{L/2-1} \sum_{q=-M/2}^{M/2-1} \sum_{r=-N/2}^{N/2-1} e^{-\frac{i\Delta t}{4} \left[ (\mu_{p}^{\widetilde{\chi}})^{2} + (\mu_{q}^{\widetilde{\chi}})^{2} + (\mu_{r}^{\widetilde{\chi}})^{2} \right]} (\widehat{\boldsymbol{\phi}_{j}^{(2)}})_{pqr} \times W_{pqr}(\widetilde{\boldsymbol{x}}_{l}, \widetilde{\boldsymbol{y}}_{k}, \widetilde{\boldsymbol{z}}_{m}).$$
(3.46)

Here,  $(\widehat{\phi_j^n})_{pqr}$  and  $(\widehat{\phi_j^{(2)}})_{pqr}$  are the discrete Fourier transform coefficients of the vectors  $\phi_j^n$  and  $\phi_j^{(2)}$ , respectively. We refer this method as TS2-GAUSUM. This scheme is explicit, efficient, unconditional



**Fig. 4.1.** The dynamics of (scaled) center of mass and trajectory for  $0 \le t \le 200$  for Cases 1–3 in Section 4.2: The first two columns for component one and the last two for component two.

stable, simple to implement and can be extended to high-order time-splitting schemes easily.

## 4. Numerical results

In this section, we first test the convergence of the TS2-GAUSUM method for computing the dynamics of rotating two-component dipolar BEC. Then, we apply our method to investigate some interesting phenomena, such as the dynamics of dipolar BEC with tunable (time-dependent) dipole axis, collapse properties of a dipolar BEC.

#### 4.1. Test of convergence

Here, we first test the spatial and temporal convergence of our method in both 2D and 3D. To demonstrate the results, we first define the following error function

$$\begin{aligned} \Psi_{\Psi}^{n,\Delta t}(t) &=: \|\Psi(\mathbf{x},t_n) - \Psi_{h,\Delta t}^{n}(\mathbf{x})\|_{l^2} \\ &= \sqrt{\sum_{j=1}^{2} \|\psi_j(\mathbf{x},t_n) - \psi_{j,h,\Delta t}^{n}(\mathbf{x})\|_{l^2}^2}, \end{aligned}$$
(4.47)

where  $\|\cdot\|_{l^2}$  denotes the discrete  $l^2$  norm,  $\psi_{j,h,\Delta t}^n$  is the numerical approximation of  $\psi_j(\mathbf{x}, t_n)$  obtained by the TS2-GAUSUM method (3.44)–(3.46) with time step  $\Delta t$  and mesh size  $h_v = h$  ( $v = \tilde{x}, \tilde{y}$ ) in 2D and  $v = \tilde{x}, \tilde{y}, \tilde{z}$  in 3D). The dipole axis **n** and interaction parameters are chosen as

$$\mathbf{n} = (1, 0, 0)^{T}, \qquad \begin{pmatrix} \beta_{11} & \beta_{12} \\ \beta_{21} & \beta_{22} \end{pmatrix} = \beta \begin{pmatrix} 1 & 0.8 \\ 0.8 & 1.2 \end{pmatrix}, \begin{pmatrix} \lambda_{11} & \lambda_{12} \\ \lambda_{21} & \lambda_{22} \end{pmatrix} = \frac{1}{20} \begin{pmatrix} \beta_{11} & \beta_{12} \\ \beta_{21} & \beta_{22} \end{pmatrix}.$$
(4.48)

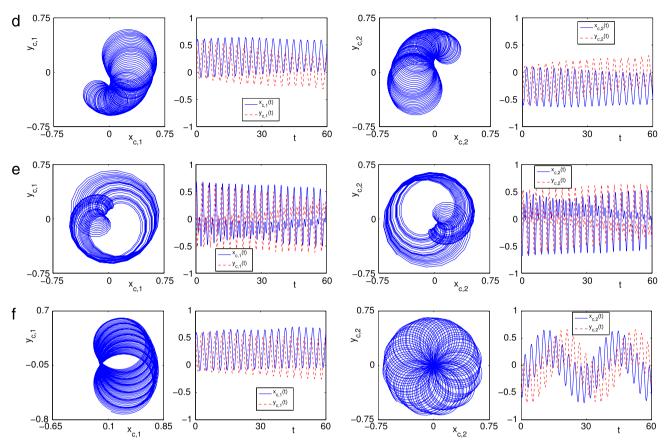
**Table 4.1** Spatial and temporal discretization errors at time t = 0.4 for the 2D CGPE with  $\Omega = 0.4$  and different  $\beta$ .

$e_{\Psi}^{h,\Delta t_0}$	h = 1	h/2	h/4	h/8
$\beta = 2$ $\beta = 10$	1.0863E-01	2.9827E-03	2.8843E-07	1.0490E-11
	3.8018E-01	4.2192E-02	7.4791E-05	1.4662E-11
$e_{\Psi}^{h_0,\Delta t}$	$\Delta t = 0.01$	$\Delta t/2$	$\Delta t/4$	$\Delta t/8$
$\beta = 2$ $\beta = 10$	2.4167E-05	6.0376E-06	1.5075E-06	3.7504E-07
	2.2051E-04	5.5049E-05	1.3742E-05	3.4187E-06

Moreover, we take the computational domain  $\mathcal{D} = [-12, 12]^2$  in 2D and  $[-8, 8]^3$  in 3D and the potential  $V_1(\mathbf{x}) = \frac{|\mathbf{x}|^2}{2}$ . The potential  $V_2(\mathbf{x})$  and initial data  $\psi_i^0(\mathbf{x})$  are chosen respectively as

$$V_{2}(\mathbf{x}) = \begin{cases} (x^{2} + y^{2})/2, \\ (x^{2} + 1.21y^{2} + z^{2})/2, \end{cases}$$
  
$$\psi_{j}^{0}(\mathbf{x}) = \begin{cases} \sqrt[4]{\frac{2}{\pi^{2}}} e^{-\frac{(3-j)x^{2}+jy^{2}}{2}}, & d = 2, \\ \sqrt[4]{\frac{2}{\pi^{3}}} e^{-\frac{(3-j)x^{2}+jy^{2}+z^{2}}{2}}, & d = 3, \end{cases}$$
  
$$j = 1, 2.$$
(4.49)

For comparison, the "exact" solution  $\Psi$  is obtained numerically via the TS2-GAUSUM method on  $\mathcal{D}$  with very small mesh size  $h = h_0 = \frac{1}{16}$  and time step  $\Delta t = \Delta t_0 = 0.0001$ . Table 4.1 lists the spatial errors  $e_{\Psi}^{h,\Delta t_0}(t)$  and temporal errors  $e_{\Psi}^{h_0,\Delta t}(t)$  at time t = 0.4 for the 2D CGPE with  $\Omega = 0.4$  and different  $\beta$ , while Table 4.2 lists those at time t = 0.1 for the 3D case with  $\Omega = 0.2$  and different  $\beta$ . From Tables 4.1–4.2, we can conclude that the TS2-GAUSUM method is second order (and spectrally) accurate in time (and space).



**Fig. 4.2.** The dynamics of (scaled) center of mass and trajectory for  $0 \le t \le 150$  for Cases 4–6 in Section 4.2: the first two columns for component one and the last two for component two.

#### Table 4.2

Spatial and temporal discretization errors at time t = 0.1 for the 3D CGPE with  $\Omega = 0.2$  and different  $\beta$ .

$e_{\Psi}^{h, \Delta t_0}$	h = 1	h/2	h/4	h/8
$\beta = 2$ $\beta = 10$	1.51E-02	1.82E-04	1.92E-08	6.60E-13
	2.60E-02	9.25E-04	8.70E-07	7.25E-13
$e_{\Psi}^{h_0,\Delta t}$	$\Delta t = 0.01$	$\Delta t/2$	$\Delta t/4$	$\Delta t/8$
$\beta = 2$ $\beta = 10$	6.14E-06	1.53E-06	3.83E-07	9.52E-08
	7.62E-05	1.90E-05	4.75E-06	1.18E-06

#### 4.2. Dynamics of the center of mass

In this subsection, we study the dynamics of the (scaled) center of mass by directly simulating the CGPE (3.29)–(3.30) via the TS2-GAUSUM method (3.44)–(3.46). To this end, we take d = 2, dipole axis  $\mathbf{n} = (1, 0, 0)^T$  and initial data (1.2)

$$\psi_1^0(\mathbf{x}) = \phi(\mathbf{x} - \mathbf{x}_0), \quad \psi_2^0(\mathbf{x}) = \phi(\mathbf{x} + \mathbf{x}_0),$$
  
with  $\phi(\mathbf{x}) = \frac{(x + iy)}{\sqrt{2\pi}} e^{-\frac{x^2 + y^2}{2}}, \quad \mathbf{x}_0 = (1, 1)^T.$  (4.50)

The computational domain, mesh size and time step are respectively taken as  $\mathcal{D} = [-16, 16]^2$ ,  $h_{\tilde{\chi}} = h_{\tilde{y}} = \frac{1}{8}$  and  $\Delta t = 0.001$ . The trapping potentials are chosen as the harmonic ones (1.3) and the following 6 cases are studied ( $j = 1, 2, k_j = 3 - j$ )

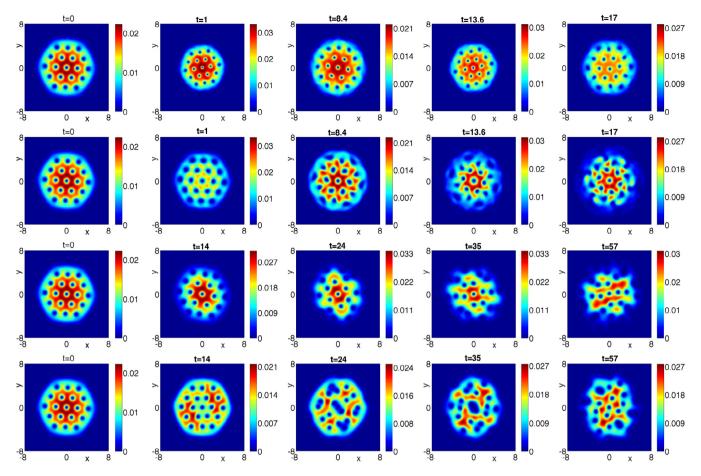
- Case 1:  $\Omega = 0.5$ ,  $\beta_{11} = \frac{\beta_{22}}{2} = 50$ ,  $\lambda_{jj} = \frac{\beta_{jj}}{10}$ ,  $\beta_{jk_j} = \frac{2\lambda_{jk_j}}{5} = 2$ ,  $\gamma_{x,j} = \gamma_{y,j} = 1$ .
- Case 2:  $\Omega = 0.5$ ,  $\beta_{jj} = 10\lambda_{jj} = 50$ ,  $\beta_{jk_j} = \frac{2\lambda_{jk_j}}{5} = 2$ ,  $\gamma_{x,1} = \gamma_{y,2} = 1.1$ ,  $\gamma_{y,1} = \gamma_{x,2} = 1$ .

- Case 3:  $\Omega = 0.5$ ,  $\beta_{jj} = 10\lambda_{jj} = 50$ ,  $\beta_{jk_j} = 2\lambda_{jk_j} = 2$ ,  $\gamma_{x,1} = \gamma_{y,1} = 1$ ,  $\gamma_{x,2} = \gamma_{y,2} = 1.2$ .
- Cases 4–6: same parameters as in Cases 1–3, except only change as  $\Omega = 1$ ,  $\Omega = \pi$  and  $\Omega = 1$ , respectively.

Figs. 4.1–4.2 show the dynamics of the (scaled) center of mass  $\mathbf{x}_{c,i}(t)$  (i = 1, 2) and its trajectory in the Cartesian coordinates for Cases 1–6. From Figs. 4.1–4.2 and additional results not shown here for brevity, we can conclude that: (i) When  $\Omega < \min\{\gamma_{x,j}, \gamma_{y,j}\},\$ then the (scaled) center of mass of the *j*th component  $\mathbf{x}_{c,i}$  always moves within a bounded domain (cf. Fig. 4.1). Otherwise, it may move helically outward (cf. Fig. 4.2). (ii) If  $V_1(\mathbf{x}) = V_2(\mathbf{x})$  with  $\gamma_{x,i} =$  $\gamma_{y,j}$ ,  $\beta_{12} = \beta_{21}$  and  $\lambda_{12} = \lambda_{21}$ , the total center of mass  $\mathbf{x}_c$  moves periodically with a period depending on both the rotating frequency and trapping frequency. In addition, the dynamics of  $\mathbf{x}_c$  does not depend on the interaction parameters  $\lambda_{ij}$  and  $\beta_{ij}$  (i, j = 1, 2), which is consistent with (3.18)–(3.19). (iii) If  $\beta_{12} = \beta_{21} \& \lambda_{12} = \lambda_{21}$ and each trapping potential is symmetric but  $V_1(\mathbf{x}) \neq V_2(\mathbf{x})$ , the interaction between two components affects the motion of  $\mathbf{x}_{c,i}$  and hence  $\mathbf{x}_c$ . Actually, the dynamical patterns of the (scaled) center of mass are similar with its single-component counterparts [47], where the center of mass always moves periodically. Nevertheless, due to the interaction between each component, the dynamics patterns are modified and  $\mathbf{x}_{c,i}$  moves now quasi-periodically (cf. Figs. 4.1(c) and 4.2(f)). (iv) If the trapping potentials are not symmetric, the dynamics of (scaled) center of mass becomes more complicated. Interactions between the two components will affect significantly the dynamics pattern of the center of mass.

## 4.3. Dynamics of quantized vortex lattices

In the following, we study the dynamics of quantized vortex lattices in the rotating two-component dipolar BECs. To this end,



**Fig. 4.3.** Contour plots of the densities  $|\psi_1(\mathbf{x}, t)|^2$  and  $|\psi_2(\mathbf{x}, t)|^2$  for Case 1 (top two rows) and Case 2 (bottom) in Section 4.3.

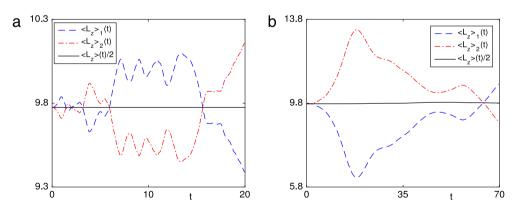


Fig. 4.4. Dynamics of the angular momentum expectation for Case 1 (a) and 2 (b) in Section 4.3.

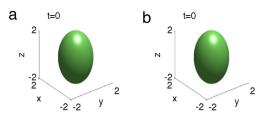
we choose d = 2,  $\beta_{11} = \beta_{22} = 100$ ,  $\beta_{12} = \beta_{21} = 70$  and  $\Omega = 0.9$ . The trapping potentials are chosen as the harmonic ones (1.3) with  $\gamma_{x,j} = \gamma_{y,j} = 1, j = 1, 2$ . The initial datum (1.2) is chosen as the stationary vortex lattice state computed by the classical gradient-flow method [41,54] for the chosen parameters without DDI, i.e.  $\lambda_{11} = \lambda_{22} = \lambda_{12} = \lambda_{21} = 0$ . The dynamics of vortex lattices are studied for the following two cases:

- Case 1: perturb the trapping frequency in component one by setting  $\gamma_{x,1} = \gamma_{y,1} = 1.5$ .
- Case 2: turn on the dipolar interaction in component one by setting  $\mathbf{n} = (1, 0, 0)^T$  and  $\lambda_{11} = 10$ .

In our simulation, we take  $\mathcal{D} = [-12, 12]^2$ ,  $h_{\tilde{\chi}} = h_{\tilde{y}} = \frac{1}{8}$  and  $\Delta t = 0.001$ . Fig. 4.3 shows the contour plots of the density

function  $|\psi_j(\mathbf{x}, t)|^2$  (j = 1, 2) at different times for Cases 1 and 2, while Fig. 4.4 shows the dynamics of the angular momentum expectation. From these two figures, we can see that: (i) The total angular momentum expectation is conserved if  $\beta_{12} = \beta_{21}$ ,  $\gamma_{x,j} = \gamma_{y,j}$  and  $\lambda_{ij} = 0$  (i, j = 1, 2), which agrees with (3.4). (ii) If there is no DDI and the trapping potentials are symmetric, the lattices rotate around the origin and keep a similar symmetry and pattern as the initial ones. Meanwhile, the lattices also undergo a breather-like dynamics. (iii) The DDI affects the dynamics very much. Due to the anisotropic nature of DDI, the lattices will rotate to some quite different patterns. The vortices will be redistributed during dynamics. Unlike the single-component BEC, the redistribution here does not seem to be aligned with the dipole axis because of the interaction between the two components. It will be interesting

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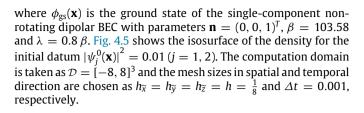
**Fig. 4.5.** Isosurface of the initial densities  $\rho_1^0(\mathbf{x}) = |\psi_1^0(\mathbf{x})|^2 = 0.01$  (a) and  $\rho_2^0(\mathbf{x}) = |\psi_2^0(\mathbf{x})|^2 = 0.01$  (b) in Example 4.1.

to investigate in further studies how the patterns of vortex lattices re-form and change with respect to the interactions as well as the dipole orientations.

#### 4.4. Numerical results in 3d

In this subsection, we report the dynamics of non-rotating twocomponent dipolar BECs in different set-ups. To this end, unless stated otherwise, the trapping potential and initial datum are chosen respectively as

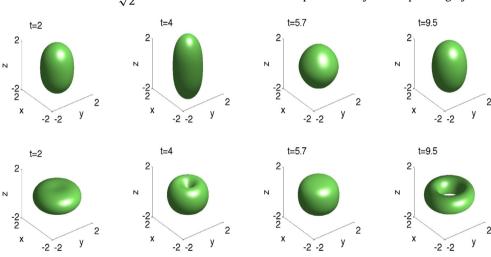
$$V_1(\mathbf{x}) = V_2(\mathbf{x}) = \frac{|\mathbf{x}|^2}{2}, \quad \psi_1^0(\mathbf{x}) = \psi_2^0(\mathbf{x}) = \frac{1}{\sqrt{2}}\phi_{gs}(\mathbf{x}), \quad (4.51)$$



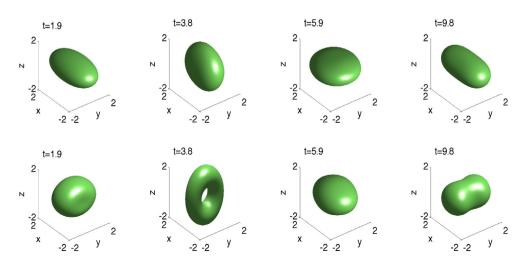
**Example 4.1.** Let  $\beta_{11} = \beta_{22} = \beta$ ,  $\lambda_{11} = \lambda$  and consider the following three cases: for  $j = 1, 2, k_j = 3 - j$ 

- Case 1: let  $\beta_{jk_j} = 100$  and turn off the DDI in component two, i.e.  $\lambda_{22} = \lambda_{jk_j} = 0$ . The dipole axis in component one is kept unchanged, i.e.  $\mathbf{n} = (0, 0, 1)^T$ .
- Case 2 : change the dipole axis to **n** = (1, 0, 0)<sup>*T*</sup> and keep the other parameters the same as in Case 1.
- Case 3: perturb the interatomic interaction as well as the DDI strength, i.e.  $\beta_{12} = \beta_{21} = 50$ ,  $\lambda_{22} = 0.8\beta$  and  $\lambda_{jk_j} = 0.8\beta_{jk_j}$ . The dipole axis is now time-dependent:  $\mathbf{n} = (\sin(t/2), 0, \cos(t/2))^T$ .

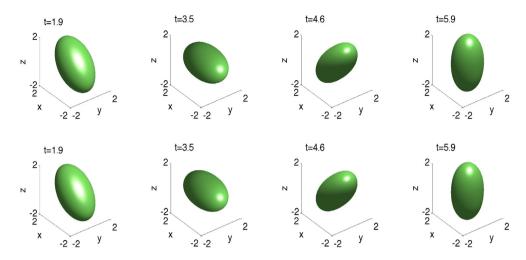
Figs. 4.6–4.8 depict the isosurface of the densities  $|\psi_j(\mathbf{x}, t)|^2 = 0.01 (j = 1, 2)$  at different times. From these figures and additional results not shown here for brevity, we can see that: (i) The total energy and mass are conserved well. (ii) Phase separation of the two components may come up during dynamics (cf. Figs. 4.6–4.7).



**Fig. 4.6.** Isosurface of the densities  $\rho_1(\mathbf{x}, t) = |\psi_1(\mathbf{x}, t)|^2 = 0.01$  (first row) and  $\rho_2(\mathbf{x}, t) = |\psi_2(\mathbf{x}, t)|^2 = 0.01$  (second row) at different times in Example 4.1: Case I.



**Fig. 4.7.** Isosurface of the densities  $\rho_1(\mathbf{x}, t) = |\psi_1(\mathbf{x}, t)|^2 = 0.01$  (first row) and  $\rho_2(\mathbf{x}, t) = |\psi_2(\mathbf{x}, t)|^2 = 0.01$  (second row) at different times in Example 4.1: Case II.



**Fig. 4.8.** Isosurface of the densities  $\rho_1(\mathbf{x}, t) = |\psi_1(\mathbf{x}, t)|^2 = 0.01$  (first row) and  $\rho_2(\mathbf{x}, t) = |\psi_2(\mathbf{x}, t)|^2 = 0.01$  (second row) at different times in Example 4.1: Case III.

In fact, the BECs would undergo mixing and de-mixing formation cyclically [15,17,22]. (iii) Similar as those shown in the singlecomponent BEC [32], when the trapping potentials are isotropic, the shapes of the density profile seem unchanged and keep the same symmetric structure with respect to the dipole orientation if the dipole axis rotates slowly (cf. Fig. 4.8).

**Example 4.2.** Here we study the collapse dynamics of the dipolar BEC. To this end, we take initial data as (4.51) with same interaction parameters and dipole axis **n** under trapping potential  $V_1(\mathbf{x}) =$  $V_2(\mathbf{x}) = \frac{x^2 + y^2 + 25z^2}{2}$ . Fig. 4.9 shows the isosurface of the densities for the initial datum  $|\psi_i^0(\mathbf{x})|^2 = 0.002$  (j = 1, 2). The computational domain and time step are chosen as  $\mathcal{D} = [-8, 8] \times [-8, 8] \times [-4, 4]$ and  $\Delta t = 0.0001$ , respectively. We consider two cases of collapse dynamics: for  $i, j = 1, 2, k_i = 3 - j$ 

- Case 1: let  $\beta_{ij} = \beta$  and change the DDI strength from  $\lambda_{ij} = \lambda$
- to  $\lambda_{11} = \lambda_{jk_j} = 2\lambda_{22} = 10\lambda$ . Case 2: let  $\lambda_{jj} = \lambda$ ,  $\lambda_{jk_j} = 0$  and change  $\beta_{ij} = \beta = 103.58$  to  $\beta_{ii} = -600.$

Figs. 4.10–4.11 depict the isosurface of the densities for  $|\psi_j(\mathbf{x}, t)|^2 = 0.002$  (j = 1, 2) at different times, while Fig. 4.12 shows the dynamics of energies. From these figures, we can see that: (i) The densities of the dipolar BECs collapse at finite time during the dynamics, i.e. the finite time blow-up of the solution is observed. This is especially clear for Case one where the contact short-range interactions are all repulsive. This reveals clearly the partial-attractive/partial-repulsive property of the DDI. (ii) The total energy and mass are conserved well before the blow-up time. They are not conserved near or after the blow-up time since the solution can no longer be resolved with a fixed mesh size and time steps.

To sum up, Examples 4.1 and 4.2 show that the dynamics of the dipolar BECs are interesting and quite complex. Different structure formations occur during dynamics and they depend heavily on the dipole orientation and the ratio between the DDI and the contact interaction strength. Moreover, the global existence and finitetime blow-up of the solution depend on the interaction parameters, which would be worth an investigation on its own.

## 5. Conclusions

We propose a robust and accurate numerical scheme to compute the dynamics of rotating two-component dipolar Bose-Einstein condensates (BEC). In rotating Lagrangian coordinates,

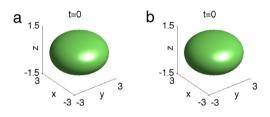
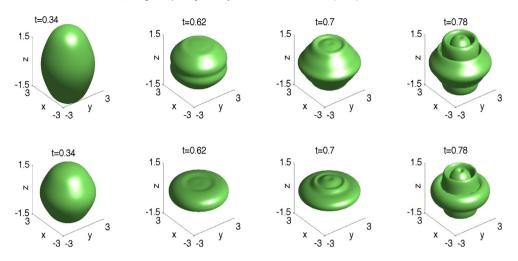


Fig. 4.9. Isosurface of the initial densities  $\rho_1^0(\mathbf{x}) = \left|\psi_1^0(\mathbf{x})\right|^2 = 0.002$  (a) and  $\rho_2^0(\mathbf{x}) = |\psi_2^0(\mathbf{x})|^2 = 0.002$  (b) in Example 4.2.

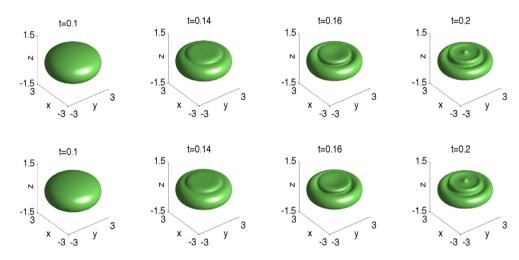
the original coupled Gross-Pitaevskii equations (CGPE) were reformulated into new equations where the rotating term vanishes. We then developed a new time splitting Fourier pseudospectral method to numerically simulate the dynamics of the new equations. The nonlocal Dipole-Dipole Interactions (DDI) were evaluated with the Gaussian-sum (GauSum) solver [51], which helps to achieve spectral accuracy within  $O(N \log N)$  operations, where N is total number of grid points. Our method is proven to be robust and efficient, and it has spectral accuracy in space and second order accuracy in time. In addition, our numerical schemes can be easily adapted and extended to tackle some more general systems, such as the spin-orbit coupled dipolar BECs [55] and the spinor dipolar BECs [8]. Further, dynamical laws of total mass, energy, center of mass and angular momentum expectation are also derived and confirmed numerically. We then applied the scheme to study the dynamics of quantized vortex lattices, the collapse dynamics of 3D dipolar BECs and identified some phenomena that are peculiar to the rotating two-component dipolar BECs.

#### Acknowledgments

We acknowledge the support by the French Science Foundation ANR under Grant No. ANR-12-MONU-0007-02 ("BECASIM") and Grant No. ANR-14-CE23-0007-01 "MOONRISE"), by the Austrian Science Foundation FWF under Grant No. F41 (SFB "VICOM"), Grant No. F65 (SFB Complexity in PDEs and Grant No. W1245 (DK "Nonlinear PDEs") and the Schrödinger Fellowship J3784-N32, by the Wiener Wissenschafts und TechnologieFonds (WWTF) Project No. MA16-066 (SEQUEX) and by the Natural Science Foundation of China Grants Nos. 11261065, 91430103 and 11471050. Further, the authors would like to acknowledge the stimulating and helpful discussions with Weizhu Bao. The computation results presented have been achieved in part by using the Vienna Scientific Cluster.



**Fig. 4.10.** Isosurface of the densities  $\rho_1(\mathbf{x}, t) = |\psi_1(\mathbf{x}, t)|^2 = 0.002$  (first row) and  $\rho_2(\mathbf{x}, t) = |\psi_2(\mathbf{x}, t)|^2 = 0.002$  (second row) at different times for collapse of Case 1 in Example 4.2.



**Fig. 4.11.** Isosurface of the densities  $\rho_1(\mathbf{x}, t) = |\psi_1(\mathbf{x}, t)|^2 = 0.002$  (first row) and  $\rho_2(\mathbf{x}, t) = |\psi_2(\mathbf{x}, t)|^2 = 0.002$  (second row) at different times for collapse of Case 2 in Example 4.2.

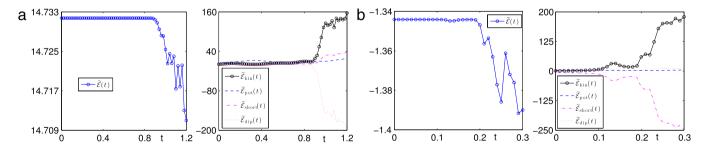


Fig. 4.12. Collapse energies for Case 1 (a) and Case 2 (b) in Example 4.2.

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